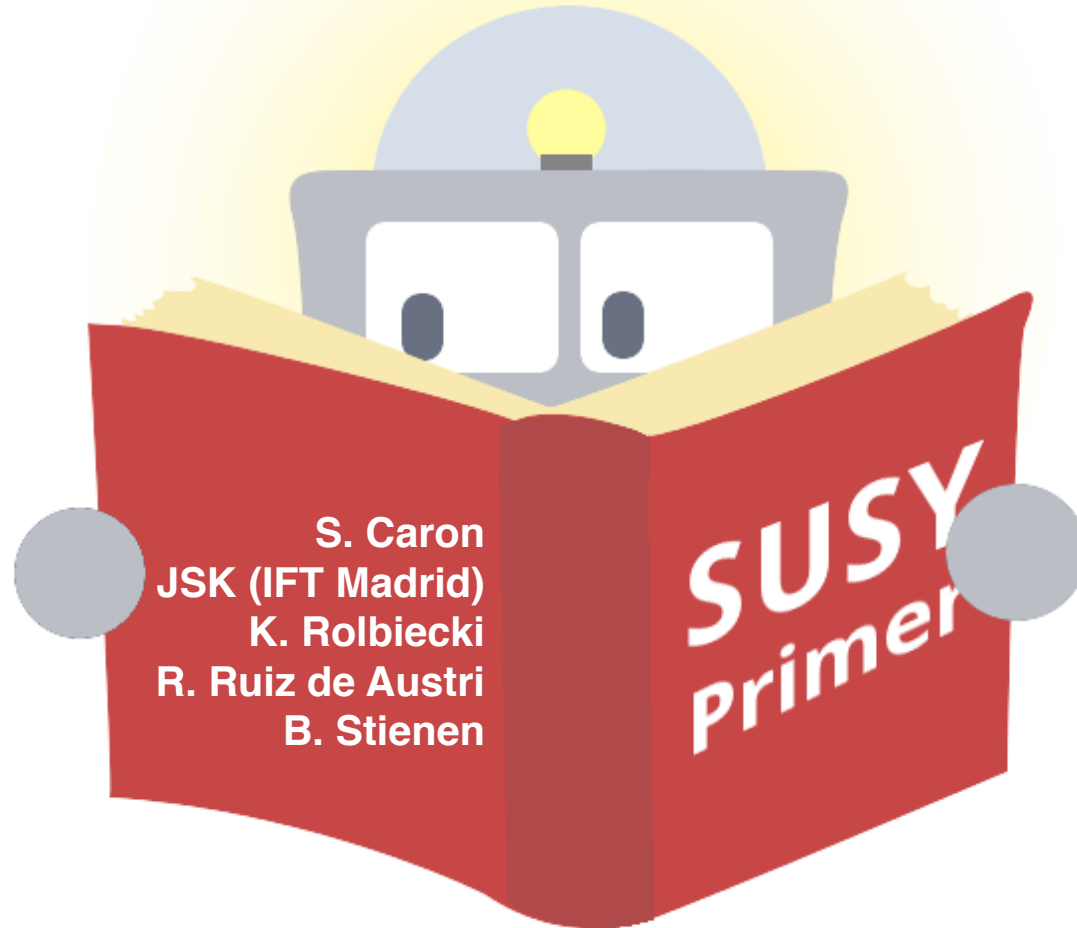


# The BSM-AI Project: SUSY-AI



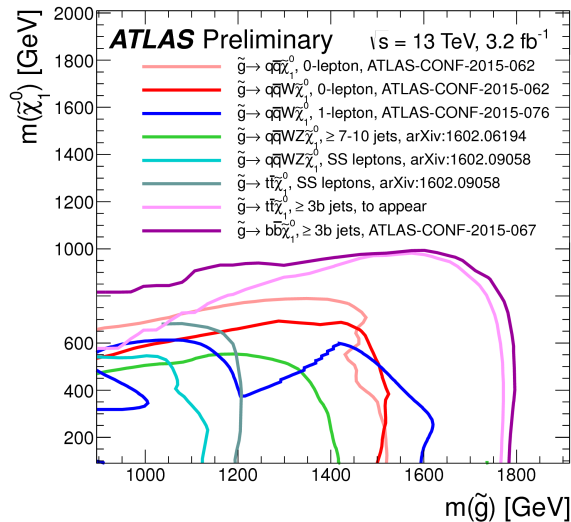
Reinterpreting the results of new physics searches at the LHC

CERN, 15-17 June 2016

Motivation

# LHC Run 2

- first Run 2 data has been analysed and no significant excess above SM expectation has been found apart from the 750 GeV resonance
- null results can be translated in strong limits on BSM scenarios



**ATLAS SUSY Searches\* - 95% CL Lower Limits**  
Status: March 2016

**ATLAS Preliminary**  
 $\sqrt{s} = 7, 8, 13 \text{ TeV}$

Model	$\epsilon, \mu, \tau, \gamma$	Jets	$L_{int}^{max}$	$[L_{int}(fb^{-1})]$	Mass limit	$\sqrt{s} = 7, 8 \text{ TeV}$	$\sqrt{s} = 10 \text{ TeV}$	Reference
<b>Inclusive Searches</b>								
MSUGRA/CMSSM	$0.3 \epsilon, \mu, 1.2 \tau$	2-10 jets/3 b	Yes	20.3	6.6	980 GeV	1.85 TeV	1507.0525
$\tilde{g}\tilde{g} \rightarrow q\bar{q}\chi_1^0$ (compressed)	$2 \epsilon, \mu$ (flat)	mono-jet	Yes	3.2	610 GeV	820 GeV	1.52 TeV	ATLAS-CONF-2015-062
$\tilde{g}\tilde{g} \rightarrow q\bar{q}W\chi_1^0$	$2 \epsilon, \mu$	2 jets	Yes	20.3	820 GeV	1.52 TeV	1.52 TeV	1503.0295
$\tilde{g}\tilde{g} \rightarrow q\bar{q}WZ\chi_1^0$	$2 \epsilon, \mu$	2 jets	Yes	3.2	1.20 TeV	1.37 TeV	1.37 TeV	ATLAS-CONF-2015-062
$\tilde{g}\tilde{g} \rightarrow q\bar{q}WZ\chi_1^0$	$2 \epsilon, \mu$	2 jets	Yes	3.0	1.20 TeV	1.37 TeV	1.37 TeV	ATLAS-CONF-2015-076
$\tilde{g}\tilde{g} \rightarrow q\bar{q}WZ\chi_1^0$	$2 \epsilon, \mu$	0-2 jets	Yes	20	1.20 TeV	1.37 TeV	1.37 TeV	1501.0555
$\tilde{g}\tilde{g} \rightarrow q\bar{q}WZ\chi_1^0$	$2 \epsilon, \mu$	0-2 jets	Yes	20	1.20 TeV	1.37 TeV	1.37 TeV	1602.06194
GMSB (V NLSIP)	$1.2 \tau + 0.1 \epsilon$	0-2 jets	Yes	3.2	1.63 TeV	1.63 TeV	1.63 TeV	1407.0502
GMSB (higgsino-bino NLSIP)	$2 \tau$	2 jets	Yes	20.3	1.34 TeV	1.34 TeV	1.34 TeV	1507.05493
GGM (higgsino-bino NLSIP)	$\gamma$	1 b	Yes	20.3	1.37 TeV	1.37 TeV	1.37 TeV	1507.05493
GGM (higgsino-bino NLSIP)	$\gamma$	2 jets	Yes	20.3	1.37 TeV	1.37 TeV	1.37 TeV	1507.05493
GGM (higgsino NLSIP)	$2 \epsilon, \mu$ (Z)	2 jets	Yes	20.3	900 GeV	900 GeV	900 GeV	1503.0290
Gravitino LSP	$0$	mono-jet	Yes	20.3	865 GeV	865 GeV	865 GeV	1502.0151
$\tilde{g}\tilde{g} \rightarrow b\bar{b}\chi_1^0$	$0$	3 b	Yes	3.3	1.78 TeV	1.78 TeV	1.78 TeV	ATLAS-CONF-2015-067
$\tilde{g}\tilde{g} \rightarrow t\bar{t}\chi_1^0$	$0$	3 b	Yes	3.3	1.78 TeV	1.78 TeV	1.78 TeV	1407.0600
$\tilde{g}\tilde{g} \rightarrow t\bar{t}\chi_1^0$	$0$	3 b	Yes	20.3	1.37 TeV	1.37 TeV	1.37 TeV	1407.0600
<b>3<sup>rd</sup> gen. squarks direct production</b>								
$\tilde{t}_1\tilde{t}_1 \rightarrow b\bar{b}\chi_1^0$	$0$	2 b	Yes	3.2	640 GeV	640 GeV	640 GeV	ATLAS-CONF-2015-066
$\tilde{t}_1\tilde{t}_1 \rightarrow b\bar{b}\chi_1^0$	$2 \epsilon, \mu$ (SS)	0-3 b	Yes	3.2	325-540 GeV	325-540 GeV	325-540 GeV	1602.0606
$\tilde{t}_1\tilde{t}_1 \rightarrow b\bar{b}\chi_1^0$	$1.2 \epsilon, \mu$	1-2 b	Yes	4.7/20.3	1117-170 GeV	200-500 GeV	200-500 GeV	1209.1102, 1407.0583
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$ or $\tilde{t}_1\tilde{t}_1 \rightarrow q\bar{q}\chi_1^0$	$0.2 \epsilon, \mu$	0-2 jets	Yes	20.3	95-198 GeV	205-719 GeV	205-719 GeV	1500.0816, ATLAS-CONF-2016-007
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$0$	mono-jet	Yes	20.3	80-245 GeV	80-245 GeV	80-245 GeV	1407.0568
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$2 \epsilon, \mu$ (natural GMSB)	1 b	Yes	20.3	190-690 GeV	190-690 GeV	190-690 GeV	1403.0232
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$3 \epsilon, \mu$ (Z)	1 b	Yes	20.3	290-610 GeV	290-610 GeV	290-610 GeV	1403.0232
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$1 \epsilon, \mu$	6 jets + 2 b	Yes	20.3	320-430 GeV	320-430 GeV	320-430 GeV	1508.0816
<b>EW direct</b>								
$\tilde{t}_1\tilde{t}_1 \rightarrow b\bar{b}\chi_1^0$	$2 \epsilon, \mu$	0	Yes	20.3	86-335 GeV	86-335 GeV	86-335 GeV	1403.0294
$\tilde{t}_1\tilde{t}_1 \rightarrow b\bar{b}\chi_1^0$	$2 \epsilon, \mu$	0	Yes	20.3	148-475 GeV	148-475 GeV	148-475 GeV	1407.0595
$\tilde{t}_1\tilde{t}_1 \rightarrow b\bar{b}\chi_1^0$	$2 \tau$	0	Yes	20.3	355 GeV	355 GeV	355 GeV	1407.0595
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$3 \epsilon, \mu$	0	Yes	20.3	423 GeV	423 GeV	423 GeV	1402.1029
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$2 \epsilon, \mu$	0-2 jets	Yes	20.3	270 GeV	270 GeV	270 GeV	1402.5294, 1402.7029
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$2 \epsilon, \mu$	0-2 jets	Yes	20.3	270 GeV	270 GeV	270 GeV	1501.0710
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$2 \epsilon, \mu$	0	Yes	20.3	423 GeV	423 GeV	423 GeV	1405.0586
GGM (higgsino NLSIP) weak prod.	$1 \epsilon, \mu + \gamma$	0	Yes	20.3	115-370 GeV	115-370 GeV	115-370 GeV	1507.05493
<b>Direct <math>\tilde{t}_1\tilde{t}_1</math> prod., long-lived <math>\tilde{t}_1</math></b>	Disapp. trk	1 jet	Yes	20.3	270 GeV	270 GeV	270 GeV	1310.0675
Direct $\tilde{t}_1\tilde{t}_1$ prod., long-lived $\tilde{t}_1$	dE/dx trk	Yes	18.4	495 GeV	495 GeV	495 GeV	495 GeV	1508.0532
Stable, stoppage $\beta$ hadron	4 $\mu$	0-1.5 jets	Yes	20.3	495 GeV	495 GeV	495 GeV	1310.0684
Measurable $\beta$ hadron	dE/dx trk	-	-	3.2	537 GeV	537 GeV	537 GeV	1409.5542
GMSB, stable $\tilde{t}_1 \rightarrow W\chi_1^0$	1 $\mu$	-	-	20.3	440 GeV	440 GeV	440 GeV	1409.5542
GMSB, $\tilde{t}_1 \rightarrow W\chi_1^0$ , long-lived $\tilde{t}_1$	2 $\mu$	-	-	20.3	1.0 TeV	1.0 TeV	1.0 TeV	1504.0162
$\tilde{g}\tilde{g} \rightarrow t\bar{t}\chi_1^0$	disapp. trk	2 jets	-	20.3	1.0 TeV	1.0 TeV	1.0 TeV	1504.0162
<b>Long-lived particles</b>								
LFV $\tilde{g}\tilde{g} \rightarrow X, Y, \nu\mu\tau/\nu\mu\tau$	$\nu\mu\tau/\nu\mu\tau$	-	-	20.3	320 GeV	320 GeV	320 GeV	1501.0703
Bilinear RPV CMSSM	$2 \epsilon, \mu$ (SS)	0-3 b	Yes	20.3	480 GeV	480 GeV	480 GeV	1404.2500
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$4 \epsilon, \mu$	0	Yes	20.3	750 GeV	750 GeV	750 GeV	1405.0586
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$3 \epsilon, \mu + \tau$	0	Yes	20.3	480 GeV	480 GeV	480 GeV	1405.0586
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$0$	0-7 jets	-	20.3	917 GeV	917 GeV	917 GeV	1502.0586
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$0$	0-7 jets	-	20.3	880 GeV	880 GeV	880 GeV	1502.0586
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$2 \epsilon, \mu$ (SS)	0-3 b	Yes	20.3	320 GeV	320 GeV	320 GeV	1404.2500
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$2 \epsilon, \mu$	2 jets + 2 b	Yes	20.3	320 GeV	320 GeV	320 GeV	1501.0703
$\tilde{t}_1\tilde{t}_1 \rightarrow W\chi_1^0$	$2 \epsilon, \mu$	2 b	-	20.3	0.4-1.0 TeV	0.4-1.0 TeV	0.4-1.0 TeV	ATLAS-CONF-2015-015
<b>Other</b>	Scalar charm, $\tilde{c} \rightarrow c\chi_1^0$	$0$	2 c	Yes	20.3	510 GeV	510 GeV	1501.0125

\*Only a selection of the available mass limits on new states or phenomena is shown.

# Constraining BSM

- experimentalists mostly interpret exclusion in simplified models assuming a few BSM particles accessible at the LHC
- in the past constrained models such as mSUGRA with its full parameter space were investigated
- the derived results are very impressive and many SUSY particles below the multi TeV scale are seemingly excluded!

# Drawbacks...

- however, many simplified models are unrealistic
- constrained models have a small number of parameters and therefore their whole parameter space can be tested against all phenomenological constraints but the conclusions are not *universal*
- even slight generalisation of the model invalidates the limits
- how can we constrain the most general realization of a model against experimental data?

# Derive my own limits...

- generate parton level events with Madgraph
- hadronize events with Pythia, Herwig, Sherpa
- calculate the NLO cross section
- simulate detector response with Delphes, PGS
- code up the relevant searches with hundreds of SRs
- validate the whole implementation

# There are no free lunches

- validation can be really painful
- a huge number of ATLAS and CMS searches are on the market
- the implementation of all those searches is time consuming
- several attempts to recast LHC limits, e.g. Atom, CheckMATE, FastLim, MadAnalysis 5, SModels and many others

# Simplified vs General

	Pro	Con
FastLim, SModels	very fast, very easy to use	only for simplified models
Atom, CheckMATE, MadAnalysis 5	are very generic	relatively slow since they depend on MC input

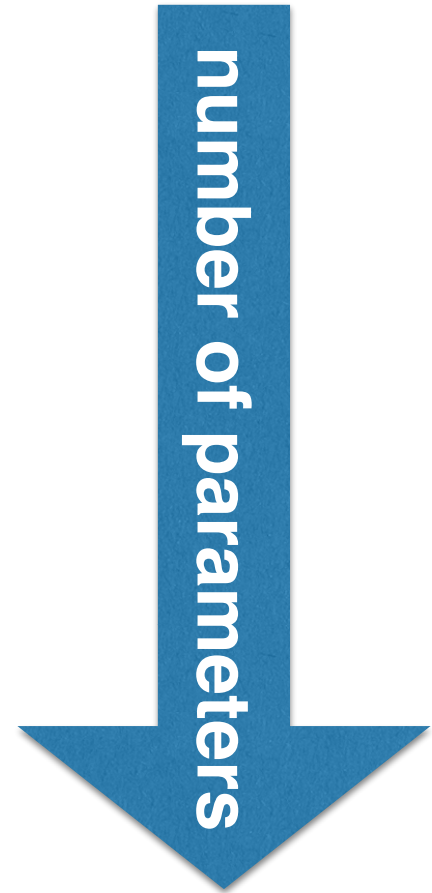


# Model complexity

simplified models  
(1-3 parameters)

constrained models, e.g. mSUGRA  
(4-6 parameters)

general models, e.g. pMSSM  
(7-20 parameters)

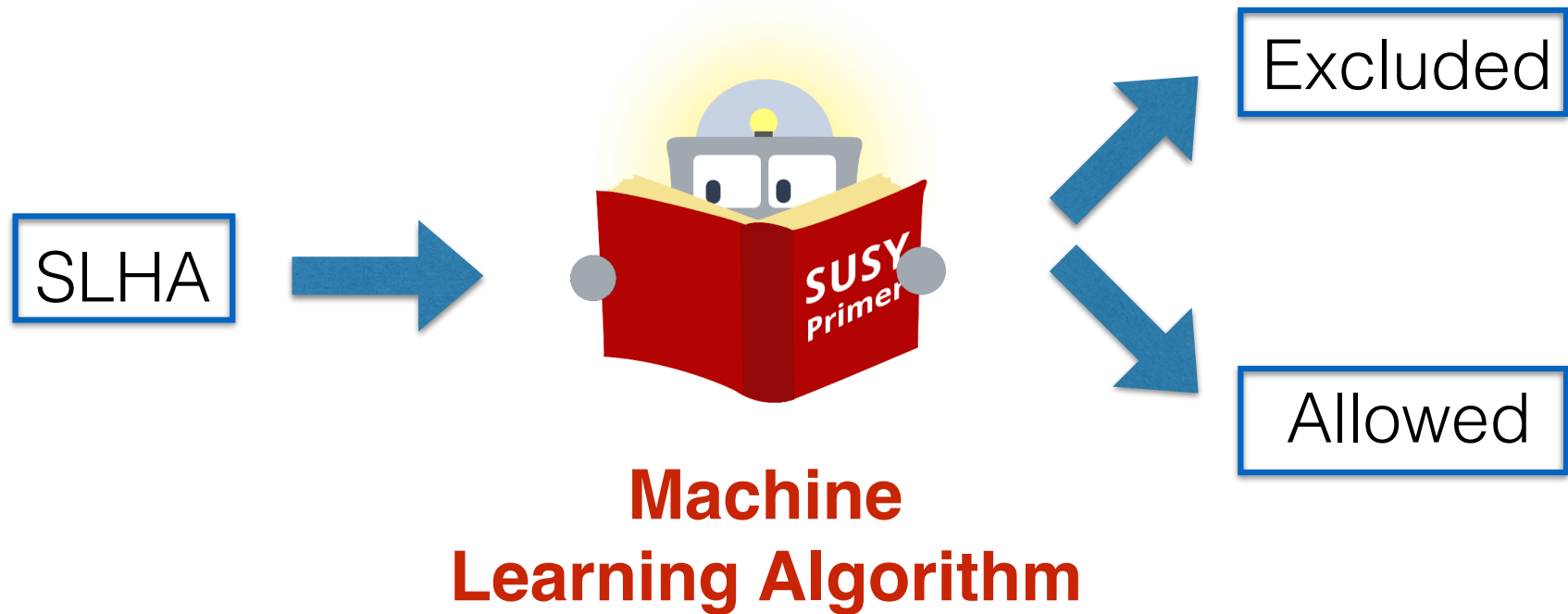


with increasing complexity, a general scan of parameter space becomes impractical (curse of dimensionality)

# Global fits

- we want to test a model against EWPO, DD and ID DM, Higgs precision measurements,  $b$  physics,...
- however, testing a complex model against collider data is difficult if the parameter space has a high dimensional volume
- e.g., in arXiv:1507.07008, we presented a new global fit of the pMSSM-19 compatible with all DM and collider constraints while accomodating the gamma ray excess from the Galactic centre
- testing the collider constraints was computationally expensive
- how can we speed up collider tests for complex models?

# Our idea: SUSY-AI



$\approx$  5000 predictions / CPU second

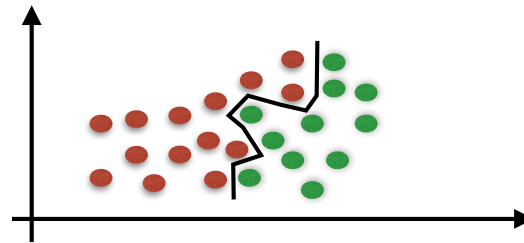
What is Machine Learning?

# Machine Learning I

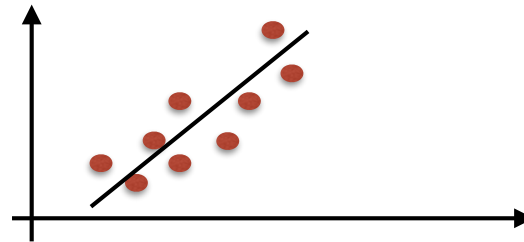
- a ML algorithm can improve its performance using training data
- the algorithm has a large number of fit parameters which can be determined by data
- ML is applied in situations which are very challenging, e.g. face or handwriting recognition

# Machine Learning II

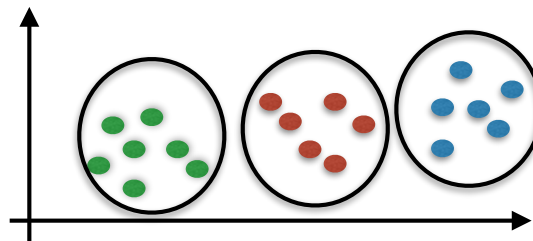
Classification - supervised



Regression - supervised



Unsupervised learning

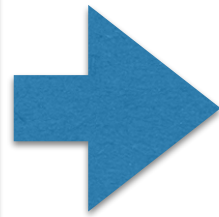


# Example

example	attribute 1	attribute 2	label
banana	...	...	1
tomato	...	...	0
cherry	...	...	1
apple	...	...	1
onion	...	...	0
cucumber	...	...	0
orange	...	...	1
water melon	...	...	?
turnip	...	...	?
maiz	...	...	?

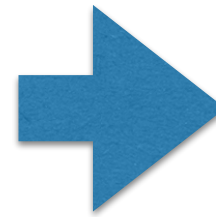
labeled (y)  
training  
data with  
attributes x

$(x,y)$



machine  
learning  
algorithm

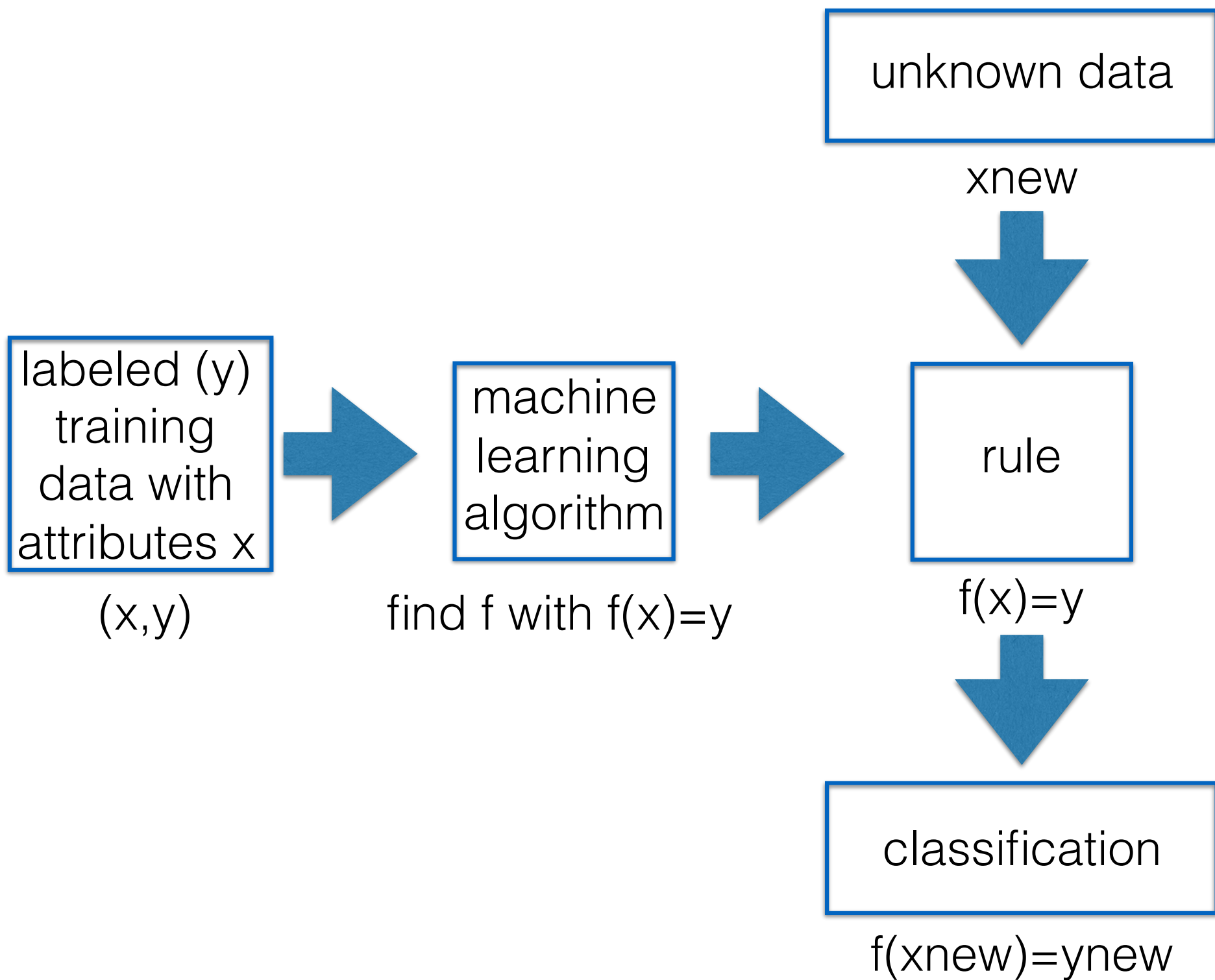
find  $f$  with  $f(x)=y$



rule

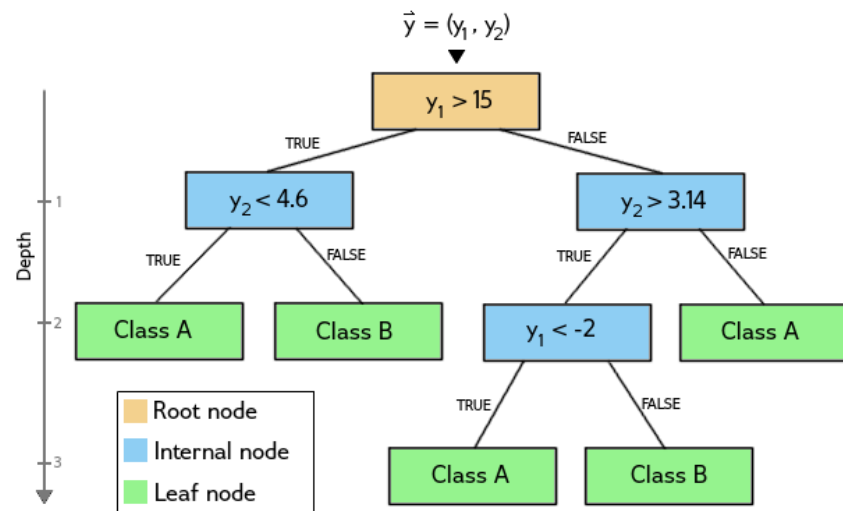
$f(x)=y$





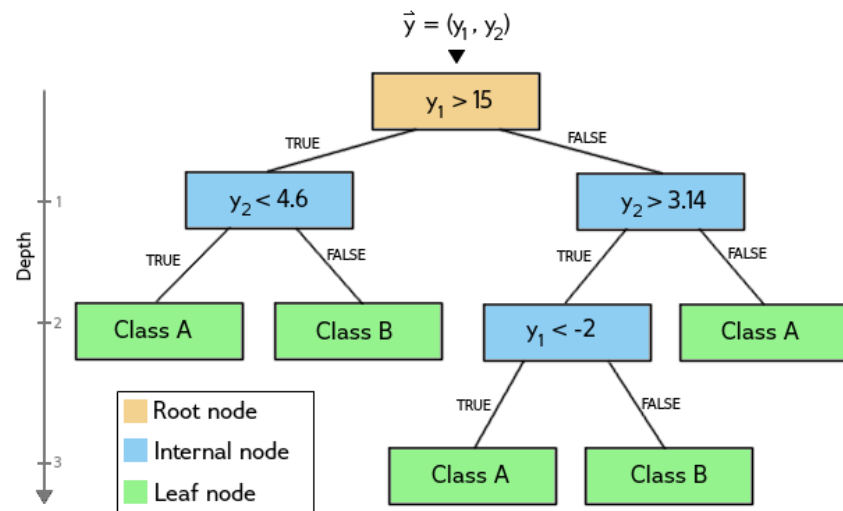
# Decision Trees I

- a Decision Tree is a commonly used classification algorithm
- DT consists of several nodes and at each node a test is performed
- the attribute set moves down the tree until the final leaf node is reached
- at the final leaf node, a class label is assigned to the attribute set



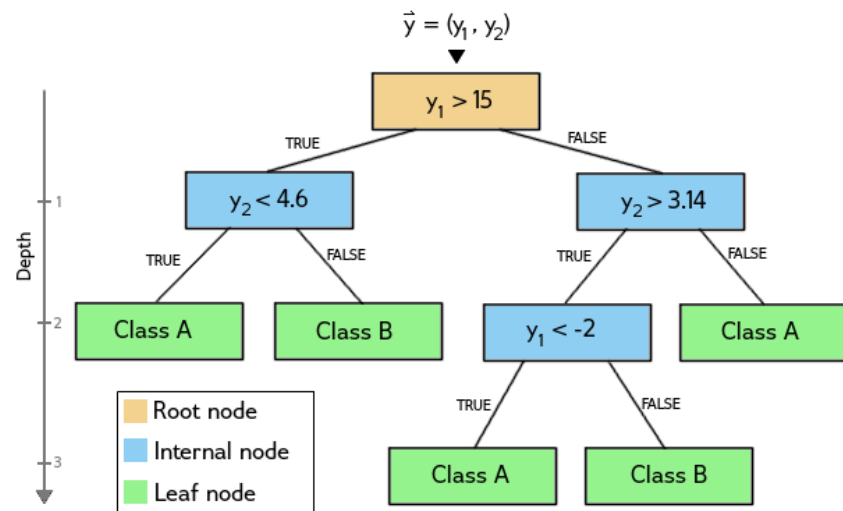
# Decision Trees II

- The DT works on the whole attribute set
- every test corresponds to a cut in this parameter space
- a DT split the attribute set into disjunct regions
- disadvantage: tendency of overtraining, i.e. DT learns the noise



# Decision Trees III

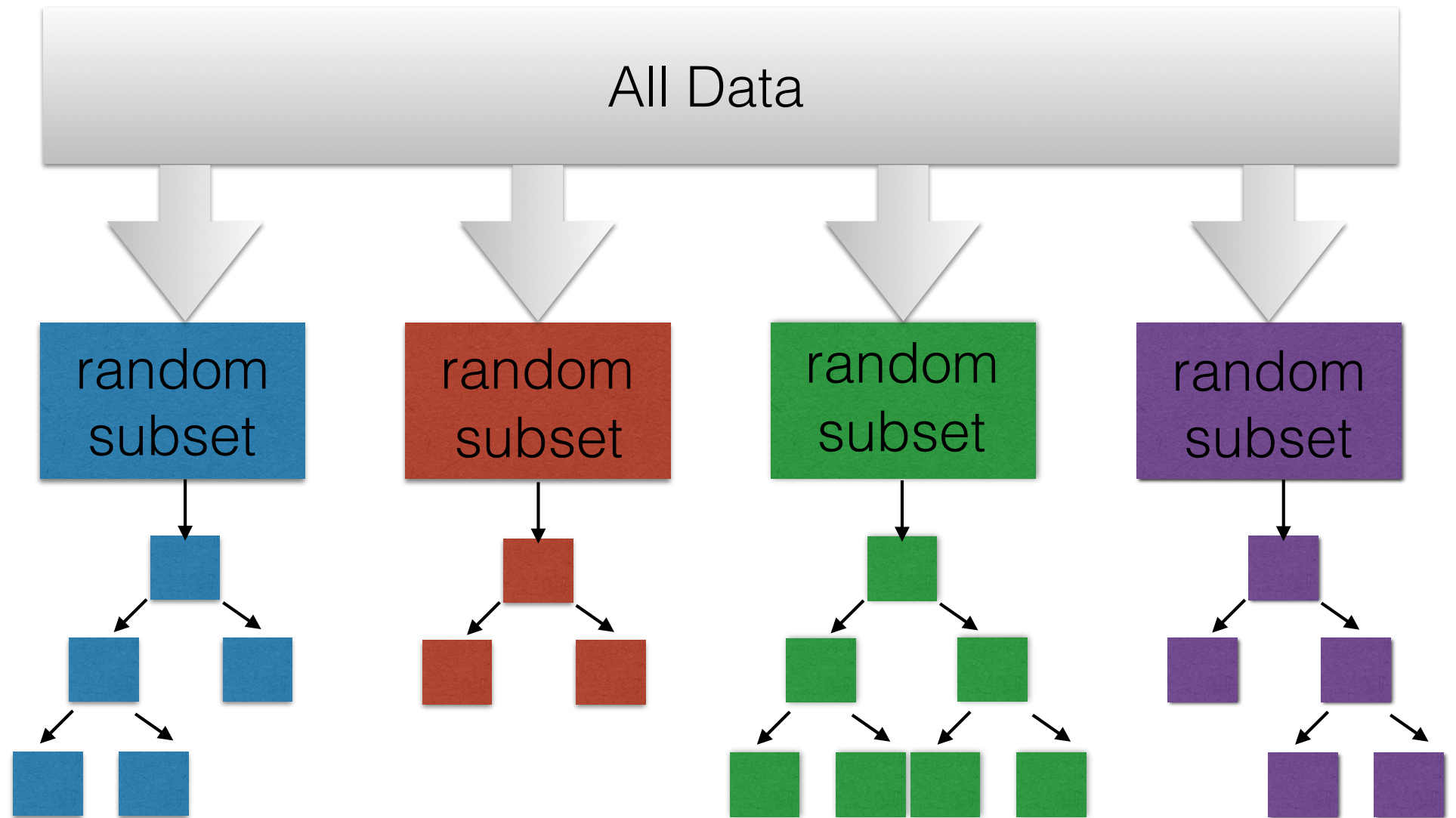
- methods to suppress overtraining
- pruning: training the entire tree but cut away all nodes beyond a certain depth
- boosting: combine multiple DTs into a single classifier
- bagging: for a group of DTs, each decision tree is trained on a random selection of a subset of the attribute set



# Random Forest I

- trees are weak learners but a forest is a strong learner
- a random forest combines trees (boosting)
- draw  $N$  bootstrap samples from original sample
- fit a classification tree to each bootstrap sample
- randomly preselect  $M$  attribute variables at each node (bagging)

# Random Forest II



at each node a random subset of the attribute set is chosen

# Random Forest III

- Output the ensemble of trees
- $R = (\text{\# trees prediction of class } C) / (\text{total number of trees})$
- $R = \text{probability of attribute set belonging to } C$
- classification in RF is done by majority vote

pMSSM-19 & ATLAS



# pMSSM-19 I

- the most general MSSM has a large number of input parameters,  $O(100)$  soft breaking parameters!
- it is unfeasible for a dedicated collider study
- assumptions on the soft breaking sector heavily reduces number of free parameters
- however, this approach might be too constraining
- consider a MSSM taking into account all constraints from particle physics experiments

# pMSSM-19 II

- consider the most general and CP conserving MSSM
- assume minimal flavour violation
- demand that the lightest neutralino is the LSP
- require the first two generation sfermions are degenerate and decoupled
- 19 weak scale parameters = pMSSM-19

# pMSSM-19 and ATLAS I

- ATLAS (arXiv:1508.06608) performed a study on the pMSSM-19
- ATLAS considered  $5 \times 10^8$  model points based on arXiv:1206.4321
- all model points had to satisfy preselection cuts
- 310,327 model points satisfy all theoretical and experimental constraints

Parameter	Description	Scanned range
$m_{\tilde{L}_1}$	1 <sup>st</sup> /2 <sup>nd</sup> gen. $SU(2)$ doublet soft breaking slepton mass	[90 GeV, 4 TeV]
$m_{\tilde{E}_1}$	1 <sup>st</sup> /2 <sup>nd</sup> gen. $SU(2)$ singlet soft breaking slepton mass	[90 GeV, 4 TeV]
$m_{\tilde{L}_3}$	3 <sup>rd</sup> gen. $SU(2)$ doublet soft breaking slepton mass	[90 GeV, 4 TeV]
$m_{\tilde{E}_3}$	3 <sup>rd</sup> gen. $SU(2)$ singlet soft breaking slepton mass	[90 GeV, 4 TeV]
$m_{\tilde{Q}_1}$	1 <sup>st</sup> /2 <sup>nd</sup> gen. $SU(2)$ doublet soft breaking squark mass	[200 GeV, 4 TeV]
$m_{\tilde{U}_1}$	1 <sup>st</sup> /2 <sup>nd</sup> gen. $SU(2)$ singlet soft breaking squark mass	[200 GeV, 4 TeV]
$m_{\tilde{D}_1}$	1 <sup>st</sup> /2 <sup>nd</sup> gen. $SU(2)$ singlet soft breaking squark mass	[200 GeV, 4 TeV]
$m_{\tilde{Q}_3}$	3 <sup>rd</sup> gen. $SU(2)$ doublet soft breaking squark mass	[100 GeV, 4 TeV]
$m_{\tilde{U}_3}$	3 <sup>rd</sup> gen. $SU(2)$ singlet soft breaking squark mass	[100 GeV, 4 TeV]
$m_{\tilde{D}_3}$	3 <sup>rd</sup> gen. $SU(2)$ singlet soft breaking squark mass	[100 GeV, 4 TeV]
$A_t$	Stop trilinear coupling	[-8 TeV, 8 TeV]
$A_b$	Sbottom trilinear coupling	[-4 TeV, 4 TeV]
$A_\tau$	Stau trilinear coupling	[-4 TeV, 4 TeV]
$ \mu $	Higgsino mass parameter	[80 GeV, 4 TeV]
$ M_1 $	Bino mass parameter	[0 TeV, 4 TeV]
$ M_2 $	Wino mass parameter	[70 GeV, 4 TeV]
$M_3$	Gluino mass parameter	[200 GeV, 4 TeV]
$M_A$	Pseudoscalar Higgs mass	[100 GeV, 4 TeV]
$\tan\beta$	Ratio of vacuum expectation values	[1, 60]

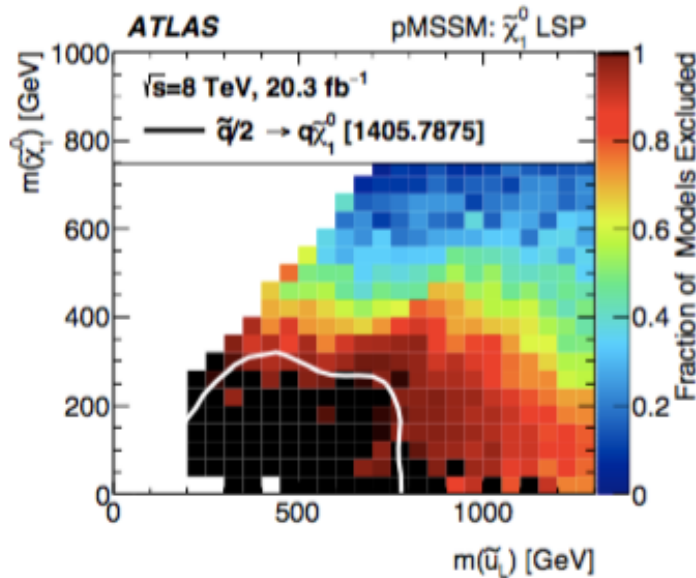
Parameter	Minimum Value	Maximum Value
$\Delta\rho$	-0.0005	0.0017
$\Delta(g-2)_\mu$	$-17.7 \times 10^{-10}$	$43.8 \times 10^{-10}$
$\text{BR}(b \rightarrow s\gamma)$	$2.69 \times 10^{-4}$	$3.87 \times 10^{-4}$
$\text{BR}(B_s \rightarrow \mu^+\mu^-)$	$1.6 \times 10^{-9}$	$4.2 \times 10^{-9}$
$\text{BR}(B^+ \rightarrow \tau^+\nu_\tau)$	$66 \times 10^{-6}$	$161 \times 10^{-6}$
$\Omega_{\tilde{\chi}_1^0} h^2$	-	0.1208
$\Gamma_{\text{invisible}}(Z)$	-	2 MeV
Masses of charged sparticle	100 GeV	-
$m_{\tilde{\chi}_1^\pm}$	103 GeV	-
$m_h$	124 GeV	128 GeV

# pMSSM-19 and ATLAS II

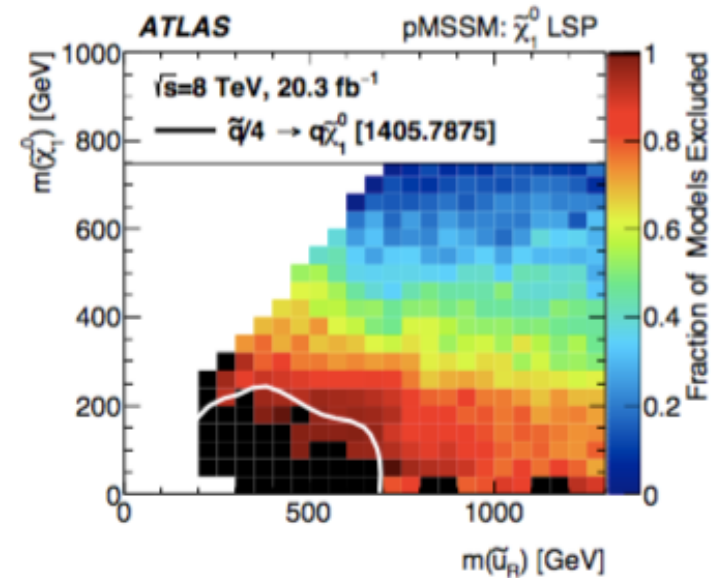
- ATLAS considered 22 separate analyses of Run 1
- a large number of final state topologies are covered
- all relevant processes were generated at truth level
- a fast detector simulation based on GEANT4 were performed

Reference	Final State	Category
[39]	0 lepton + 2 – 6 jets + $\cancel{E}_T$	Inclusive
[40]	0 lepton + 7 – 10 jets + $\cancel{E}_T$	
[41]	1 lepton + jets + $\cancel{E}_T$	
[42]	$\tau(\tau/\ell)$ + jets + $\cancel{E}_T$	
[43]	SS/3 lepton + jets + $\cancel{E}_T$	
[44]	$b$ jets + 0/1 lepton + $\cancel{E}_T$	
[45]	monojet	
[46]	0 lepton stop search	Third generation squarks
[47]	1 lepton stop search	
[48]	2 lepton stop search	
[49]	monojet search	
[50]	stop search with $Z$ in final state	
[51]	$2b$ jet sbottom search	
[4]	asymmetric stop search	
[52]	1 lepton plus Higgs final state	Electroweak
[53]	dilepton final state	
[54]	$2\tau$ final state	
[55]	trilepton final state	
[56]	four-lepton final state	
[57]	disappearing track	
[58, 59]	Long-lived particle search	Other
[60]	$H/A \rightarrow \tau\tau$ search	

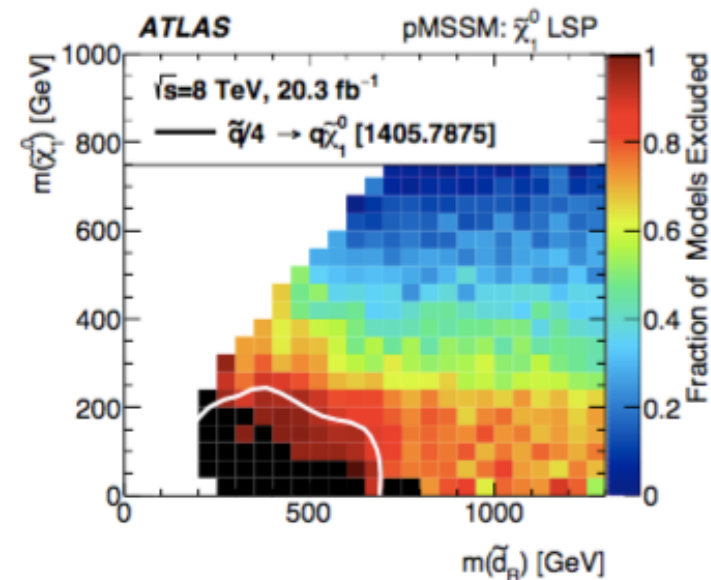
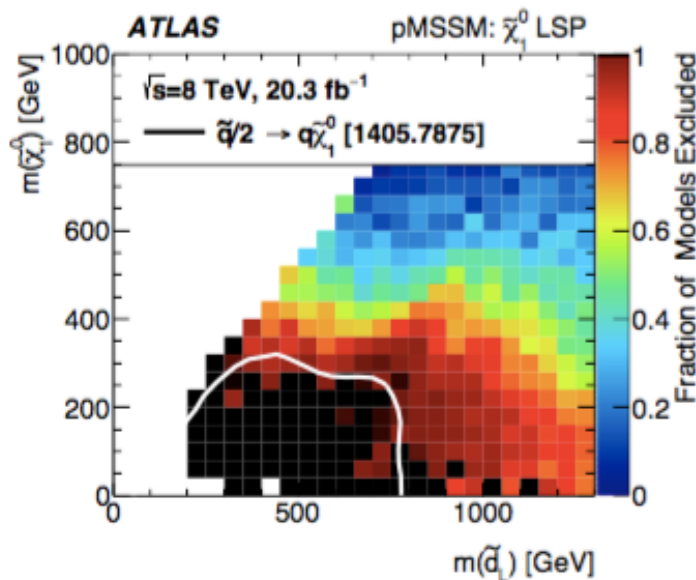
# pMSSM-19 and ATLAS III



(a) Left up squark



(b) Right up squark



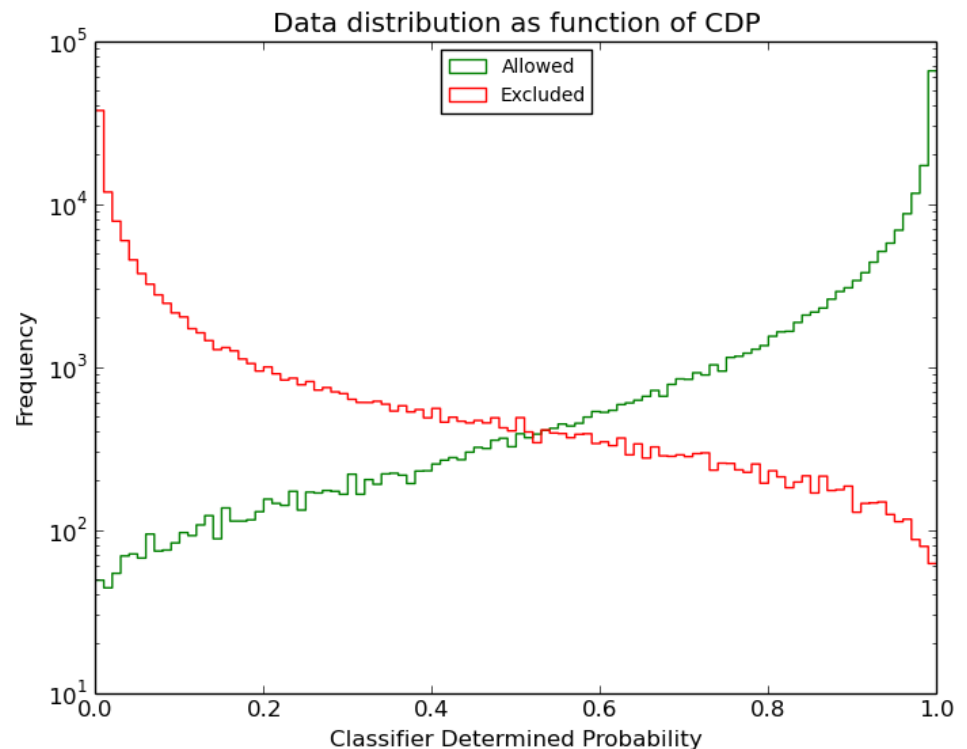
Training

# Training of SUSY-AI I

- we used the Python package *scikit-learn-0.17.1*
- we trained our RF classifier with the ATLAS data points
- we determined the optimal classifier configuration in a grid search
- 900 DT with a maximal depth of 30 nodes and a maximum number of features considered at each node of 12

# Training of SUSY-AI II

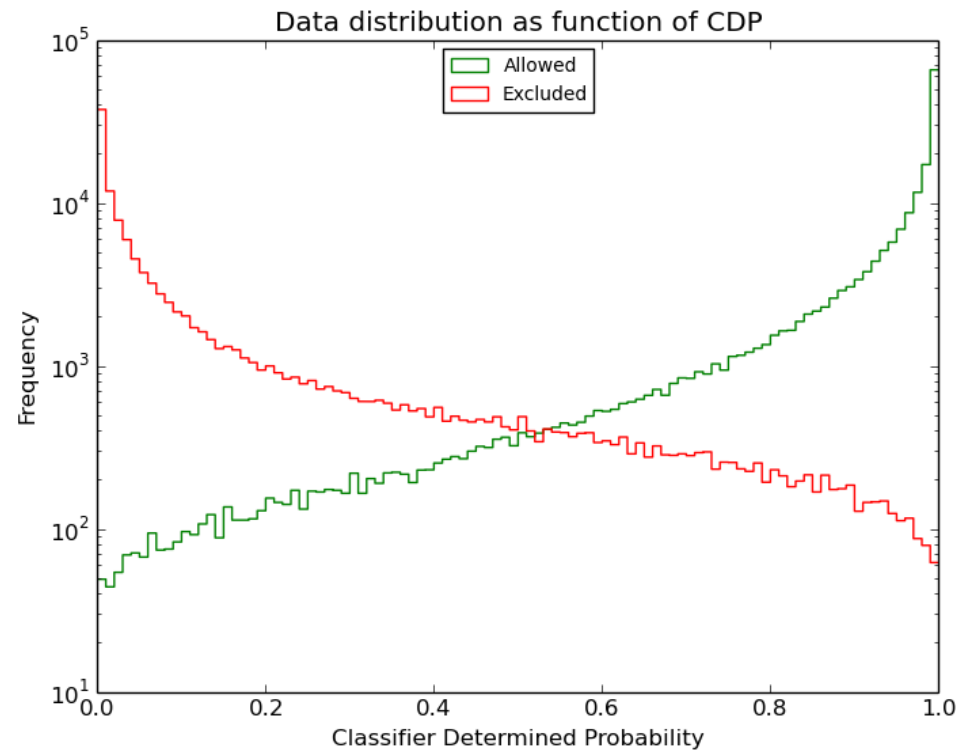
- all predicted data points are assigned with a classification probability by the RF classifier
- the green histogram includes all points which are truly allowed
- the red histogram includes all truly excluded points
- the x-axis corresponds to the classifier determined probability (CDP)





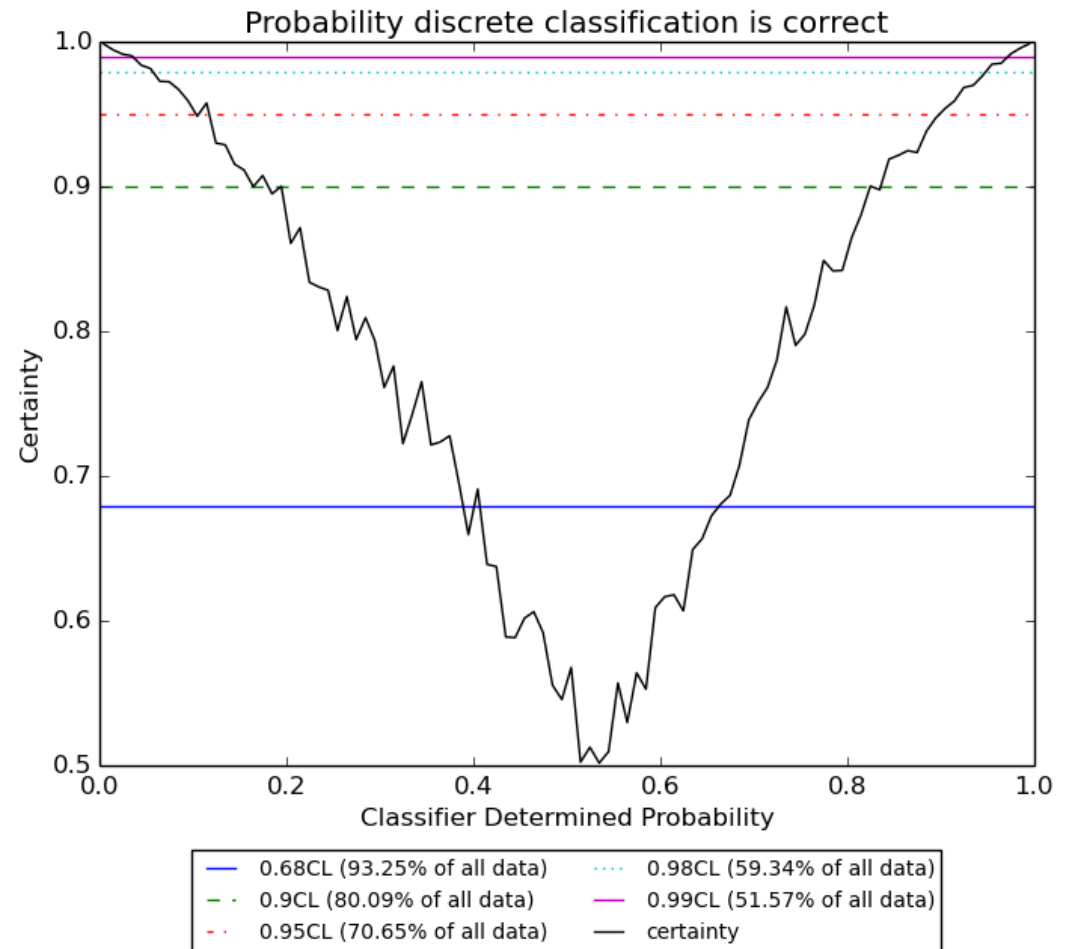
# Training of SUSY-AI III

- CDP is the probability that the model point is allowed
- majority of points are correctly classified
- however, perfect classification is not possible
- a cut makes the classification binary, e.g. a cut at 0.5, i.e. for  $\geq 0.5$ , point is allowed



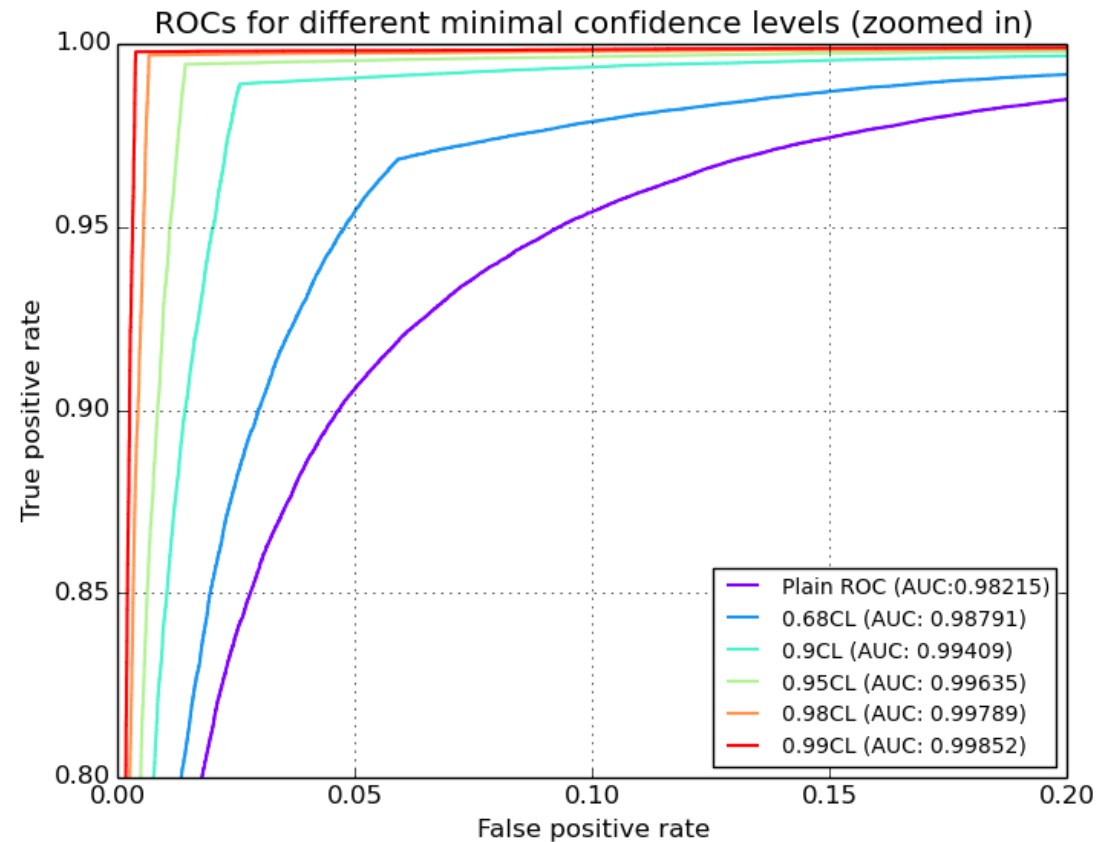
# Training of SUSY-AI IV

- take ratio of upper histogram and total number of points in each bin
- it allows a frequentist confidence level that a point with given CDP is allowed or excluded
- e.g.: a CL of 98% corresponds to a CDP of below 0.05 or above 0.95
- a CL of 95% corresponds to predicted probabilities below 0.133 or above 0.9



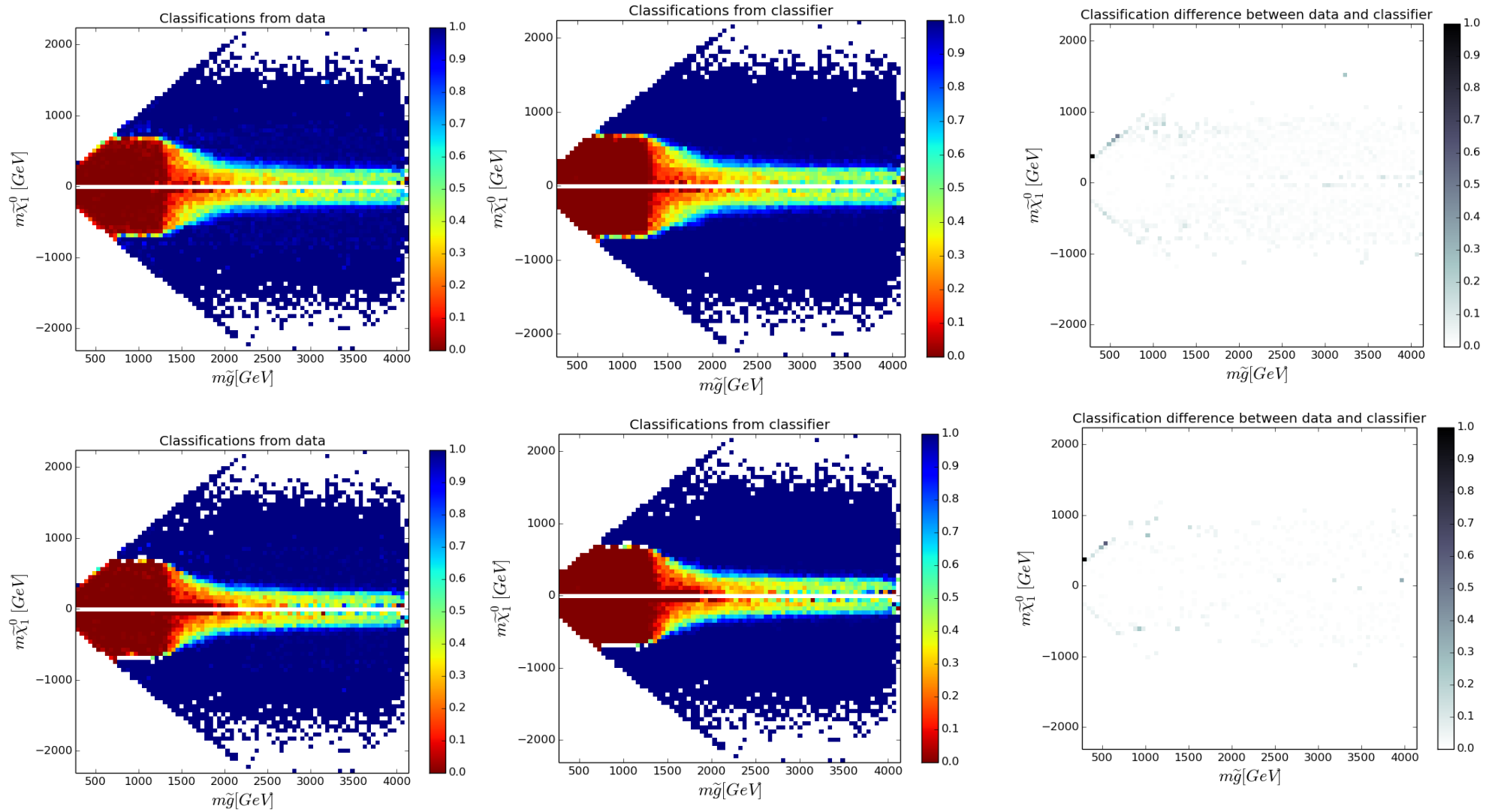
# Training of SUSY-AI V

- a “harder” cut provides more reliable results for classification but larger number of points are removed
- the performance can be quantified by the ROC curve
- higher CL cuts increases the AUC

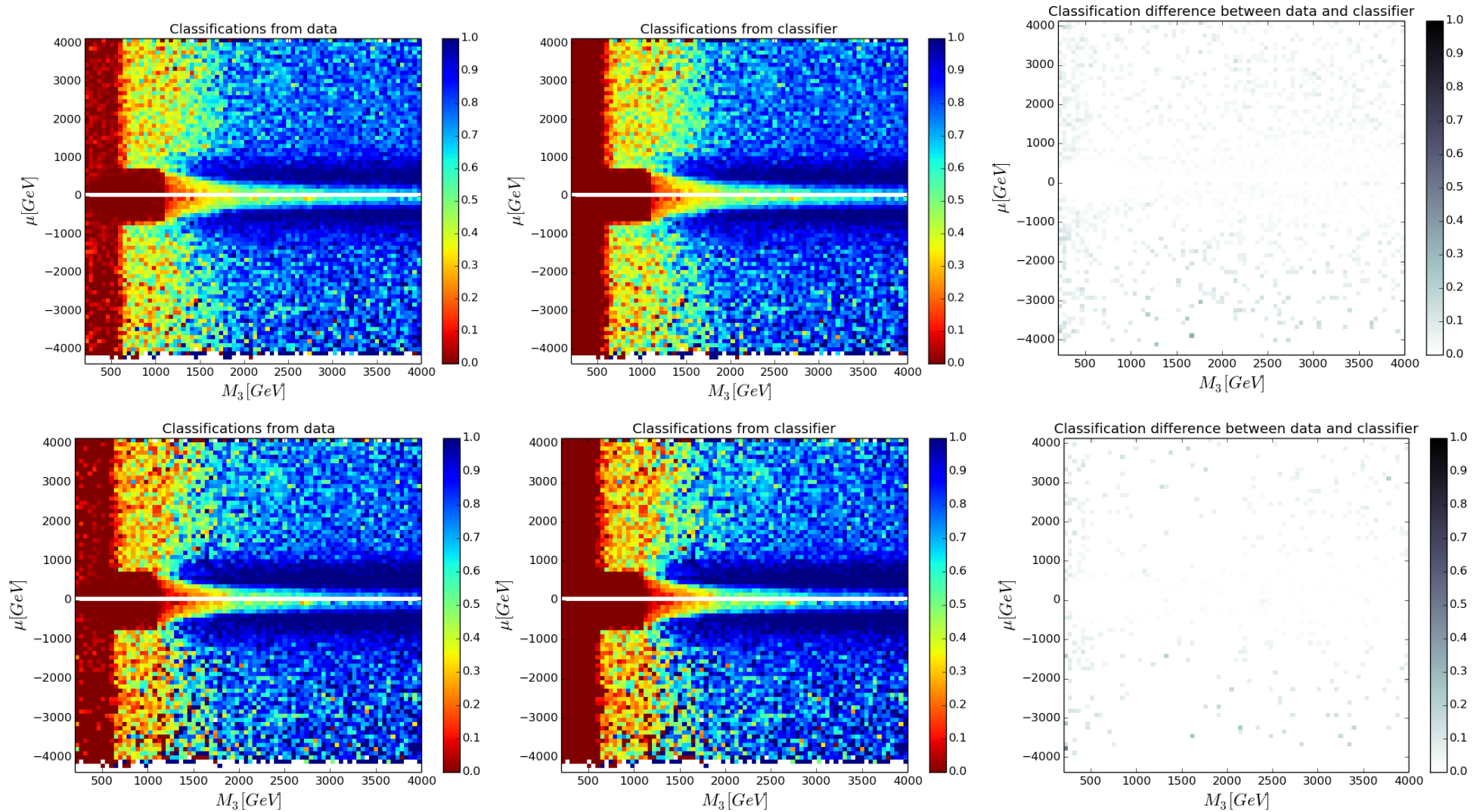


Performance

# Performance of SUSY-AI I



# Performance of SUSY-AI II



# Applications

# natural SUSY I

- a minimal natural SUSY scenario consists of light higgsinos, SU(2) doublet third generations squarks, a SU(2) singlet stop and multi-TeV gluinos
- the scenario consists of six input soft breaking parameters
- 22000 benchmark points were generated and the produced MC events analysed with CheckMATE
- we derived limits in minimal NSUSY parameter space (arXiv:1511.04461)

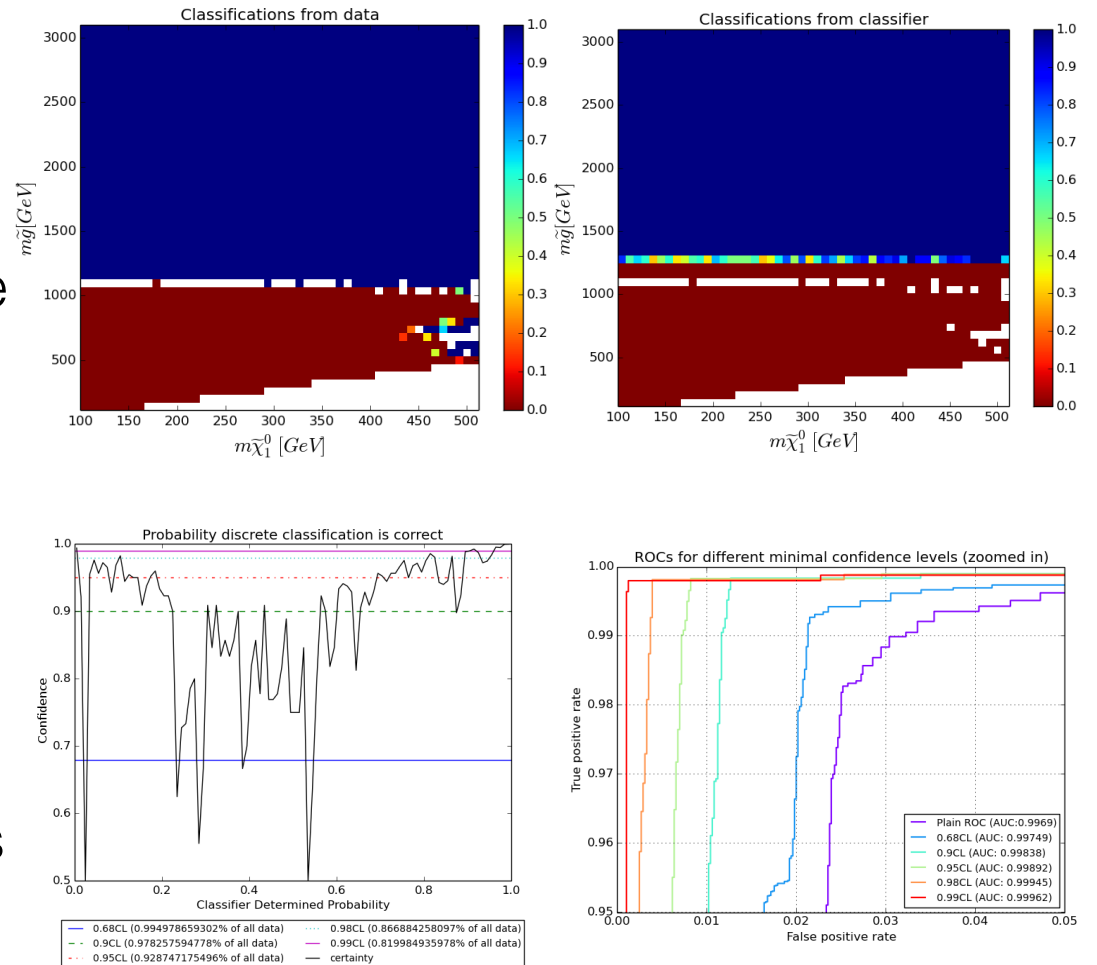
Parameter	Description	Scanned range
$m_{\tilde{Q}_3}$	3 <sup>rd</sup> generation SU(2) doublet soft breaking squark mass	[0.1 TeV, 1.5 TeV]
$m_{\tilde{U}_3}$	3 <sup>rd</sup> generation SU(2) singlet soft breaking squark mass	[0.1 TeV, 1.5 TeV]
$M_3$	Gluino mass parameter	[0.1 TeV, 3.0 TeV]
$A_t$	Stop trilinear coupling	[-3.0 TeV, 3.0 TeV]
$\mu$	Higgsino mass parameter	[0.1 TeV, 0.5 TeV]
$\tan\beta$	Ratio of vacuum expectation values	[1, 20]

Reference	Final State	$\mathcal{L}$ [fb <sup>-1</sup> ]	#SR
1308.2631 (ATLAS) [51]	0 $\ell$ +2 $b$ jets+ $\cancel{E}_T$	20.1	6
1403.4853 (ATLAS) [48]	2 $\ell$ + $\cancel{E}_T$	20.3	12
1404.2500 (ATLAS) [43]	SS 2 $\ell$ or 3 $\ell$	20.3	5
1407.0583 (ATLAS) [47]	1 $\ell$ +( $b$ ) jets+ $\cancel{E}_T$	20.0	27
1407.0608 (ATLAS) [49]	monojet+ $\cancel{E}_T$	20.3	3
1303.2985 (CMS) [89]	$\alpha_T$ + $b$ jets	11.7	59
ATLAS-CONF-2012-104 [90]	1 $\ell$ + $\geq 4$ jets+ $\cancel{E}_T$	5.8	2
ATLAS-CONF-2013-024 [91]	0 $\ell$ +6 (2 $b$ ) jets+ $\cancel{E}_T$	20.5	3
ATLAS-CONF-2013-047 [92]	0 $\ell$ +2-6 jets+ $\cancel{E}_T$	20.3	10
ATLAS-CONF-2013-061 [93]	0-1 $\ell$ + $\geq 3b$ jets+ $\cancel{E}_T$	20.1	9
ATLAS-CONF-2013-062 [94]	1-2 $\ell$ +3-6 jets+ $\cancel{E}_T$	20.0	19
CMS-SUS-13-016 [95]	OS 2 $\ell$ + $\geq 3b$ jets	19.7	1



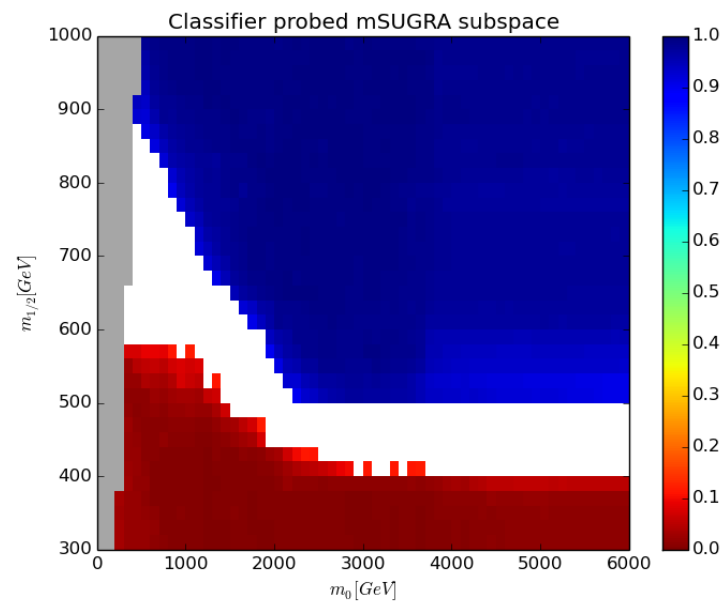
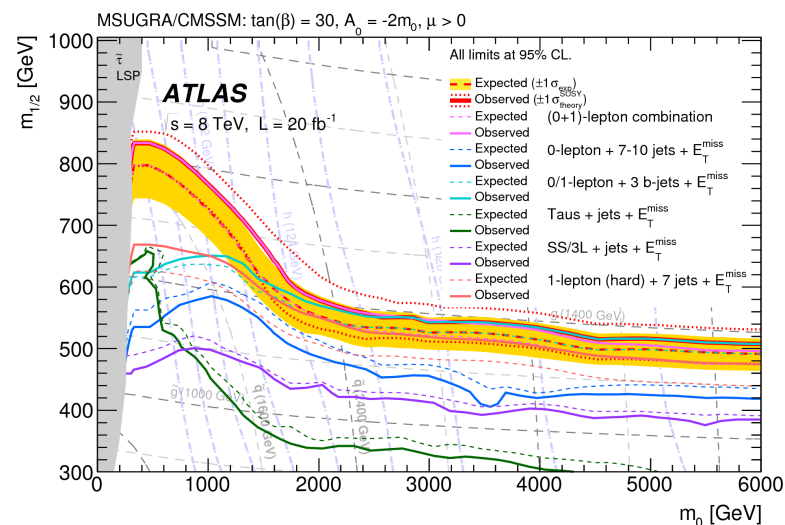
# natural SUSY II

- we tested all 22k benchmark points with SUSY-AI
- we were able to reproduce the limits
- we derived somewhat better results since the procedure with CheckMATE was conservative
- there are wrong classifications but a confidence level cut provides reliable results



# mSUGRA

- we performed a test with the constrained SUGRA model
- it has 4 and 1/2 parameters:
- $m_0$ ,  $m_{1/2}$ ,  $A_0$ ,  $\tan \beta$  and  $\text{sign } \mu$
- we set  $\tan \beta=0$  and  $A_0=2m_{1/2}$
- all points outside of the sampling range were relocated into the sampling region



The Tool

# The Tool

```
from susyai import susyai
import numpy as np

sa = susyai("susyai_classifier_python_v3.pkl")
data = np.array([[30, 4.0276e2, 7.3196e2, 2.1862e3, 1.0,
                  4.0713e3, 4.4890e3, 4.4752e3, 4.4743e3, 2.8806e3,
                  3.7855e3, 1.3240e3, 2.9076e3, 4.2226e3, 4.2056e3,
                  3.4290e3, 3.8608e3, -4.3154e3, -8.1538e3, -7.3680e3]])
clas, pred, cert = sa.predict(data)
```

```
from susyai import susyai
import numpy as np

sa = susyai("susyai_classifier_python_v3.pkl")
sa.set_coordinate_selector(1)
sa.set_id_selector('filename')

files = ['spectrum1.slha', 'spectrum2.slha']
clas, pred, cert, coords, ids = sa.predict_files(files)
```

Outlook

# BSM-AI and SUSY-AI

- we will provide classifiers for the MSSM and the NMSSM updated with 13 TeV data based on a larger training set
- we want to perform the difficult task of predicting the efficiencies/likelihoods (interesting for people performing a global fit)
- we want to include non collider constraints
- we work on providing classifiers for non-SUSY models
- ultimate goal is to consider a generic model independent approach

# Conclusion

- we trained a RF classifier on over 310,000 data points of the pMSSM-19
- we used the results from the ATLAS (arXiv:1508.06608) pMSSM study
- we obtain the correct classification with an accuracy of at least 93.8%
- we will continuously update SUSY-AI with future LHC results
- we want to provide classifiers for other BSM

<http://susyai.hepforge.org>